

Off-axis Neutrinos

Gina Rameika

Fermilab

May 20, 2006

Presentation to NuSAG
May 20, 2006

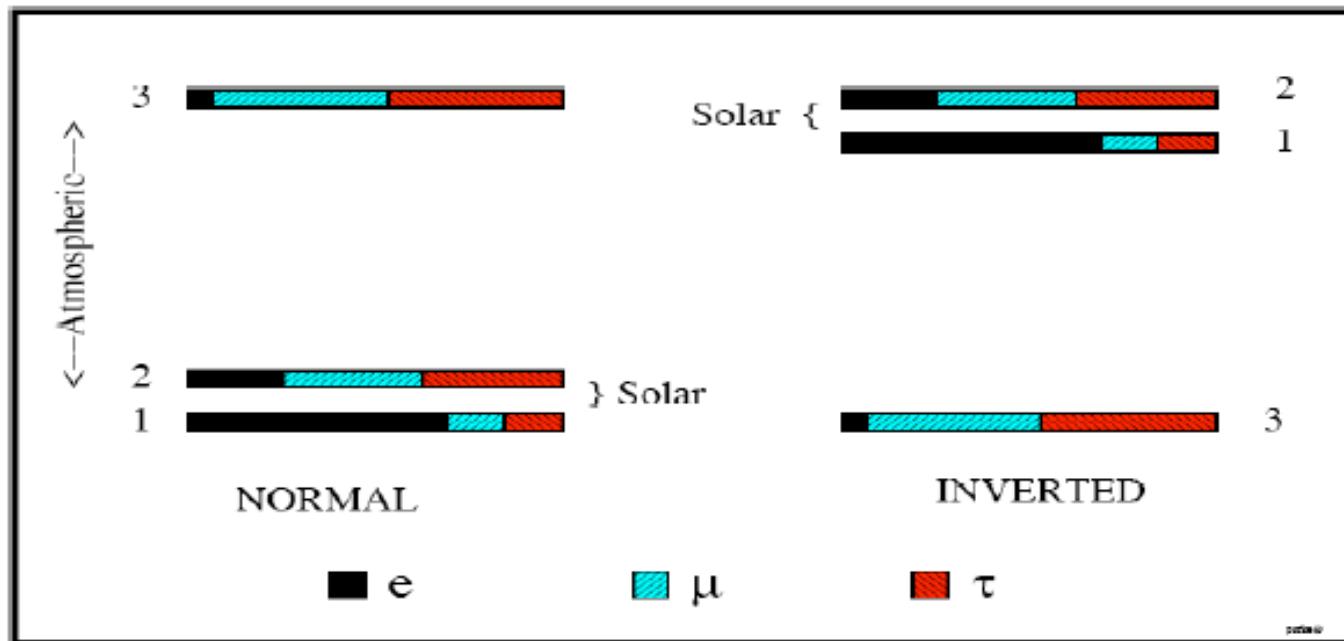
Outline

- Introduction
 - The Questions to Answer
 - The Equations
 - The Experimental Approach
- The Off-Axis Technique
 - The Basic Concept
 - The Experiments we can do
 - The Measurements we can make
 - The Physics we can extract
 - NuMI Off-Axis
 - Application in NOvA
 - Application for a NOvA2
- On-going Work - input to the LBL study
- Conclusions

The Questions

(in order of increasing difficulty)

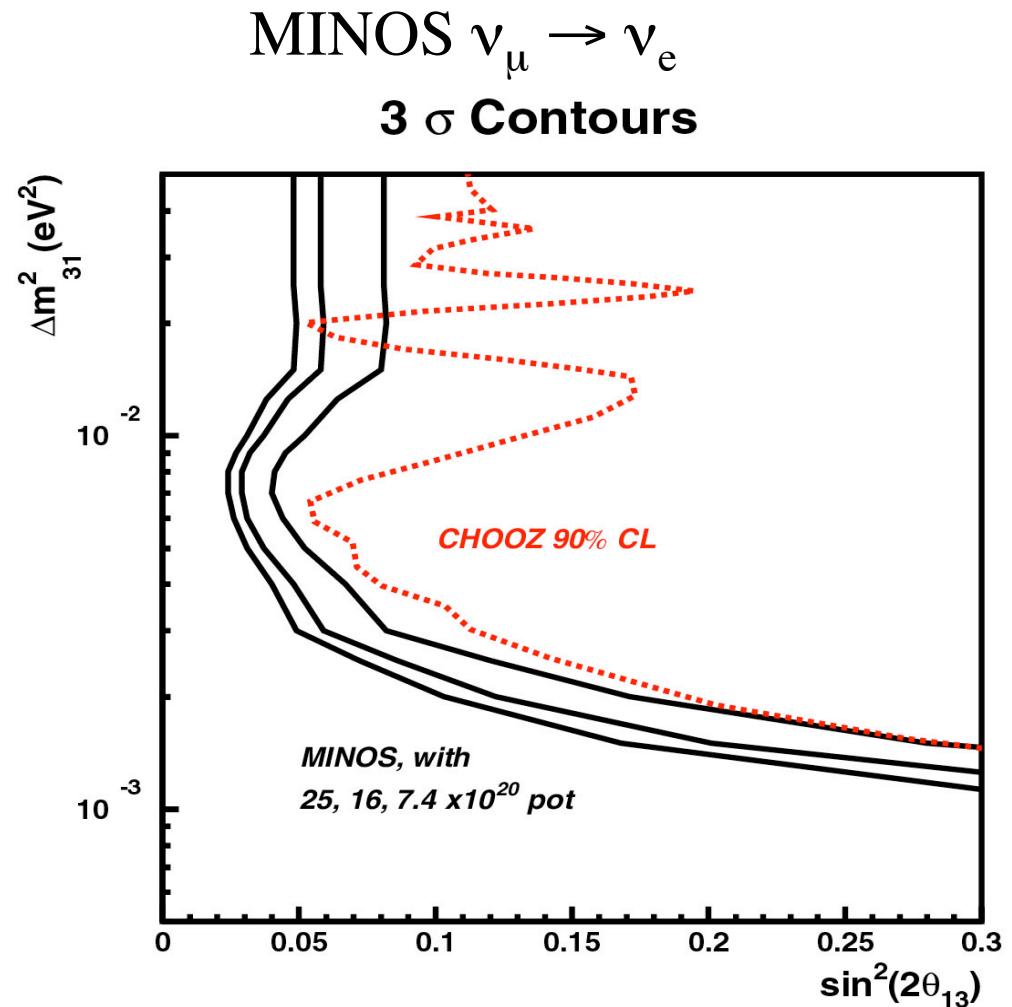
- What is $\sin^2\theta_{13}$?
- What is the order of the neutrino mass hierarchy?
- Is CP violated in the neutrino sector?
 - i.e. is $\delta \neq 0$?



Presentation to NuSAG
May 20, 2006

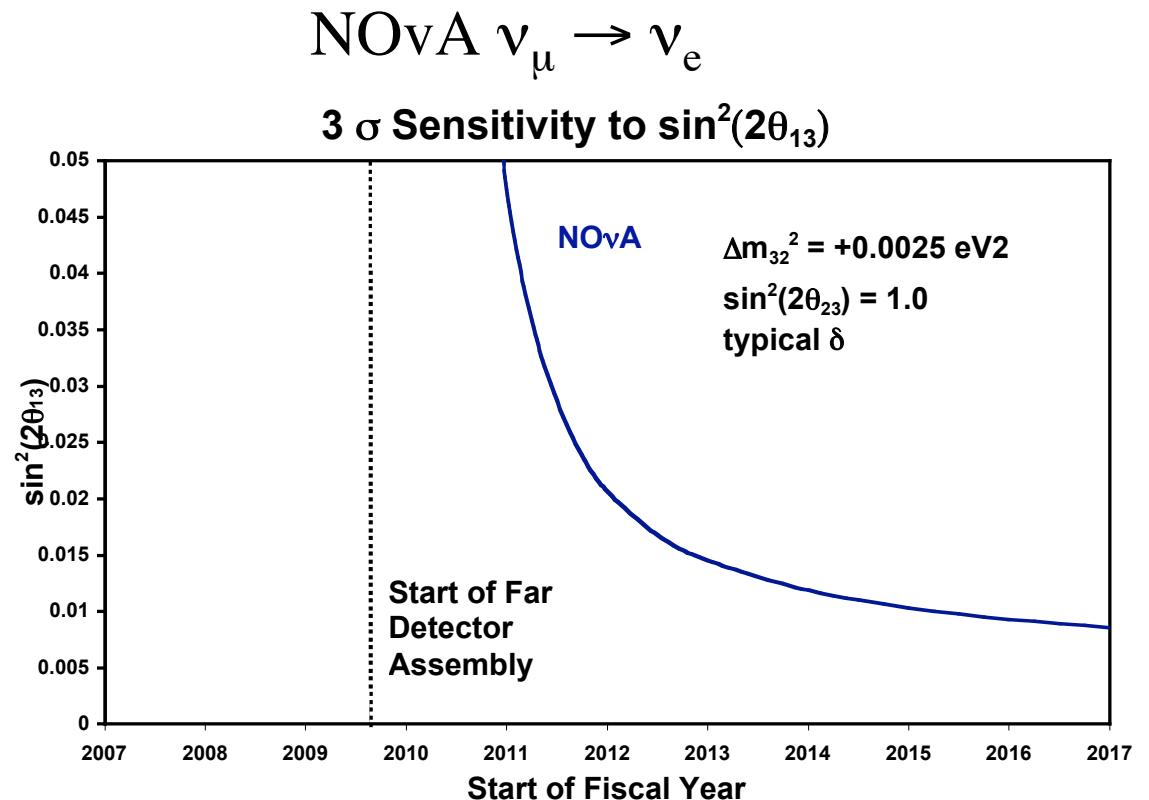
The Answers : a staged approached

- Step 1 : Current Program
 - operations 2005 - 2009-10
 - NuMI to MINOS
 - L = 735 km on-axis
 - LE beam
 - 12e20 POT
- Step 2 : Proposed Program
 - 2011 - 2016
 - NuMI to NOvA
- Step 3 : Current Discussion
 - A decade from now
 - Where we start, depends on outcome of Steps 1, 2 and other worldwide efforts

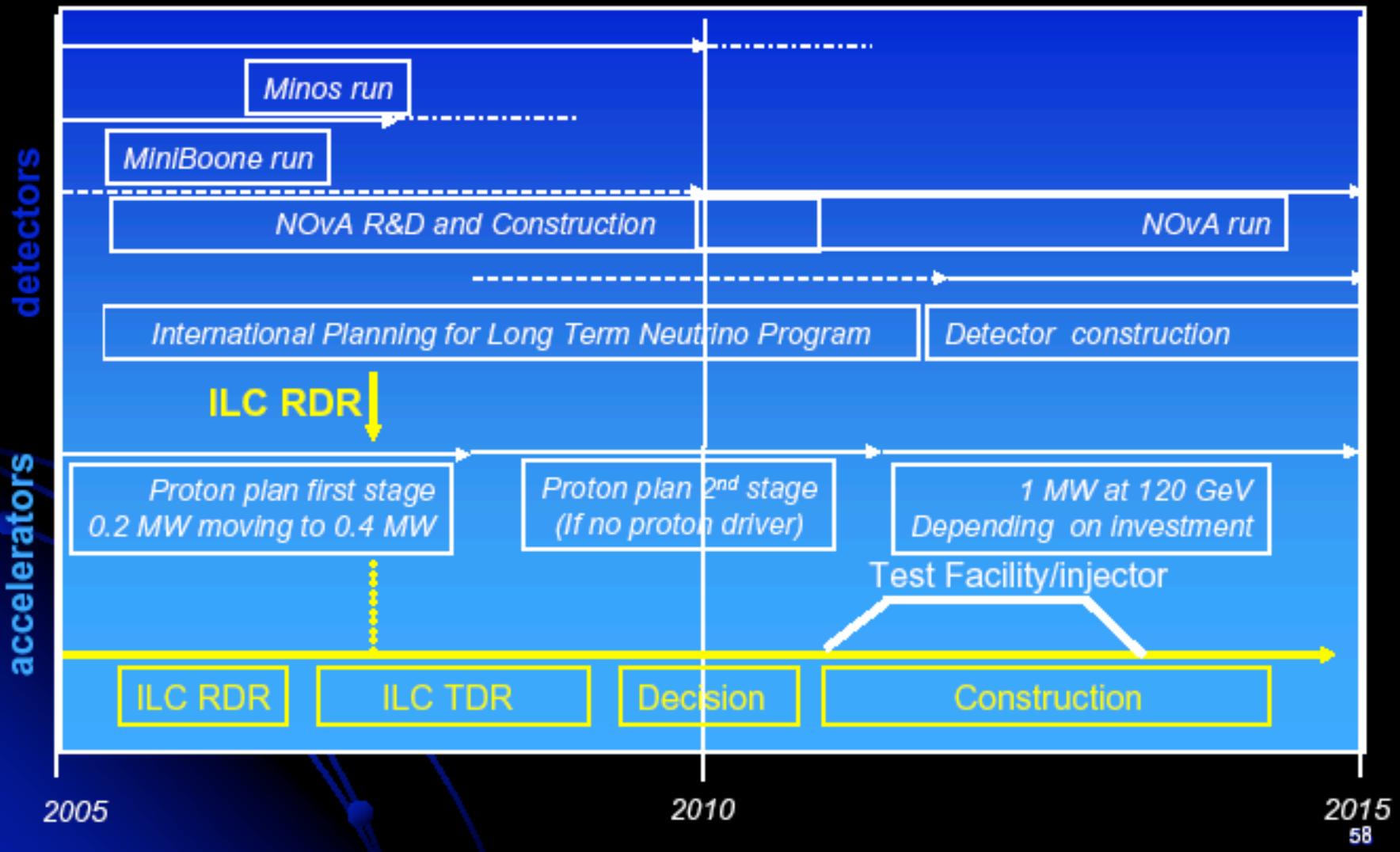


The Answers : a staged approached

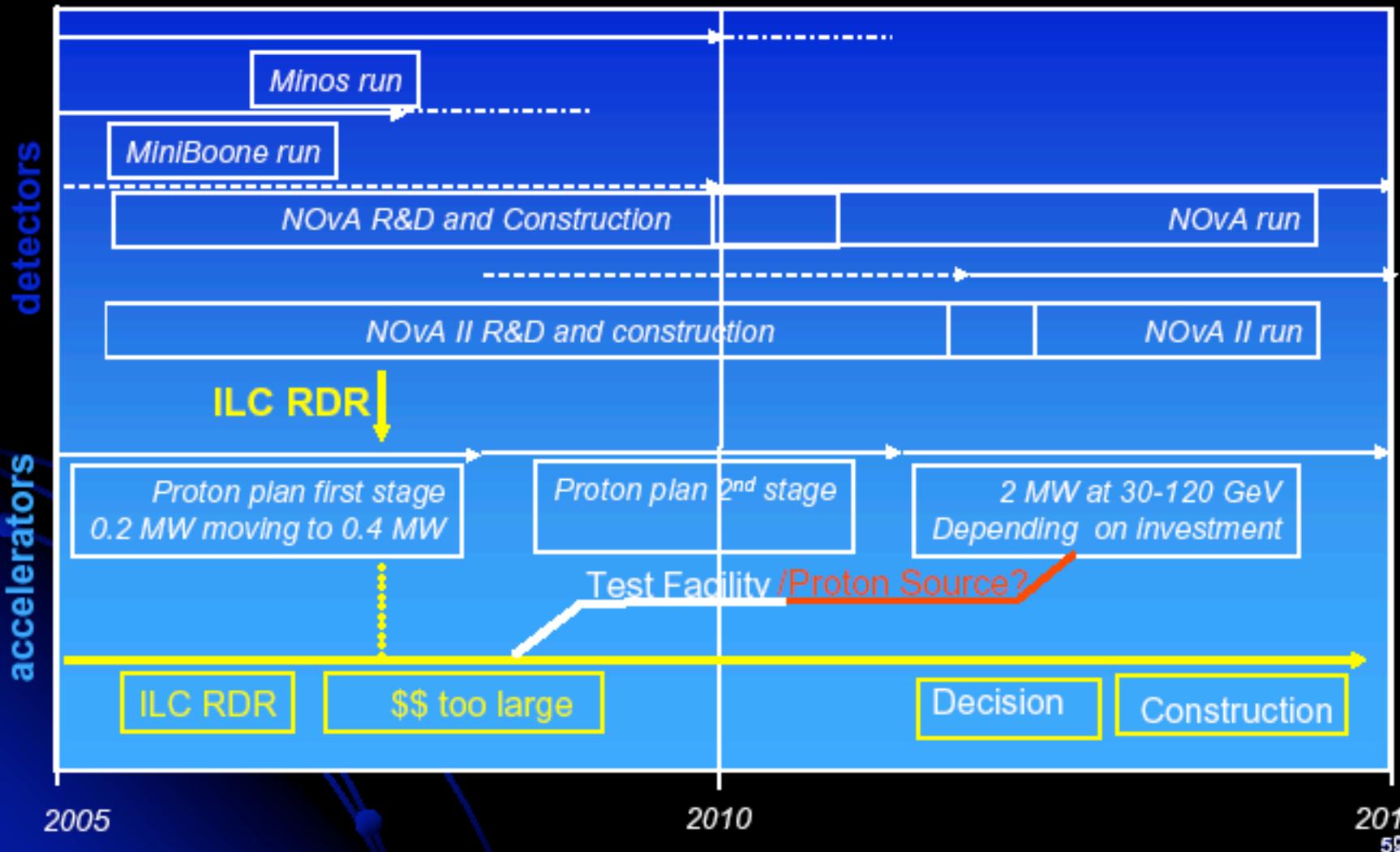
- Step 1 : Current Program
 - operations 2005 - 2009-10
 - NuMI to MINOS
 - L = 735 km on-axis
 - LE beam
 - 12e20 POT
- Step 2 : Proposed Program
 - 2011 - 2016
 - NuMI to NOvA
- Step 3 : Current Discussion
 - A decade from now
 - Where we start, depends on outcome of Steps 1, 2 and other worldwide efforts



Accelerator Programs



Accelerator Programs



Neutrino Mixing Matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & \text{solar} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \text{atmospheric}$$

The Equations

- The simple version :

$$P(\nu_a \rightarrow \nu_b) = \sin^2 2\theta_{ab} \sin^2(1.27 \Delta m_{ab}^2 L / E)$$

- Works for the dominant oscillation mode
 - $\nu_\mu \rightarrow \nu_\tau$ Atmospheric
- Extend the simple version to the subdominant mode :

$$P_{vac} (\nu_\mu \rightarrow \nu_e) = \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \Delta_{atm},$$

$$\Delta_{atm} \approx 1.27 \left(\frac{\Delta m_{32}^2 L}{E} \right),$$

The Equations continue

- Matter effects

$$P_{mat}(\nu_\mu \rightarrow \nu_e) \approx \left(1 \pm 2 \frac{E}{E_R}\right) P_{vac}(\nu_\mu \rightarrow \nu_e)$$

$$E_R = \frac{\Delta m_{32}^2}{2\sqrt{2}G_N} = 12 \text{ GeV} \left(\frac{\Delta m_{32}^2}{2.5 \times 10^{-3} \text{ eV}^2} \right) \left(\frac{1.4 \text{ g cm}^{-3}}{\text{Y}_\odot} \right)$$

$$\Delta P_\delta(\nu_\mu \rightarrow \nu_e) \approx J_r \sin \Delta_{sol} \sin \Delta_{atm} (\cos \delta \cos \Delta_{atm} + \sin \delta \sin \Delta_{atm}),$$

- CP phase....

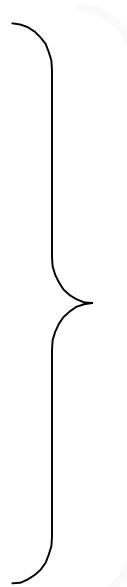
$$\Delta P_\delta(\nu_\mu \rightarrow \nu_e) \approx J_r \sin \Delta_{sol} \sin \Delta_{atm} (\cos \delta \cos \Delta_{atm} + \sin \delta \sin \Delta_{atm}),$$

$$J_r = \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13},$$

The Experimental Approach

- Depends on the “initial conditions”...
 - Existing accelerator ?
 - Proton/neutrino energies
 - Existing beamline ?
 - Beam direction
 - Desirable site ?
 - Baseline
 - Underground?
 - On surface?
 - Existing Detector ?
 - Detection threshold/efficiency
 - Baseline

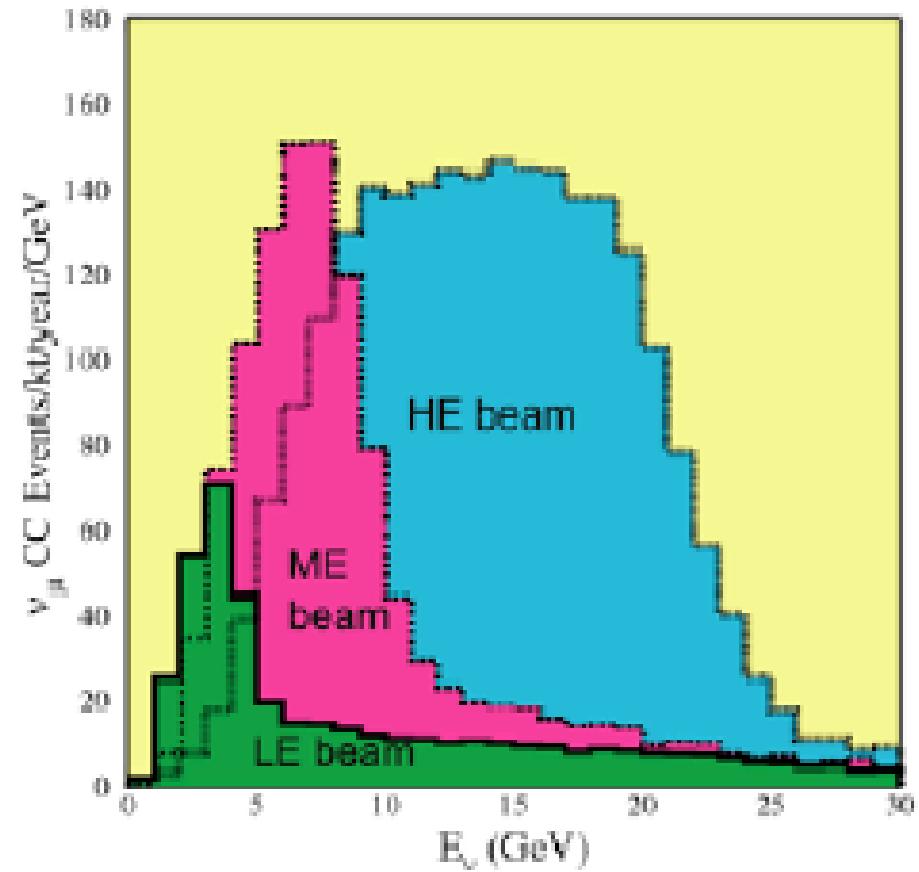
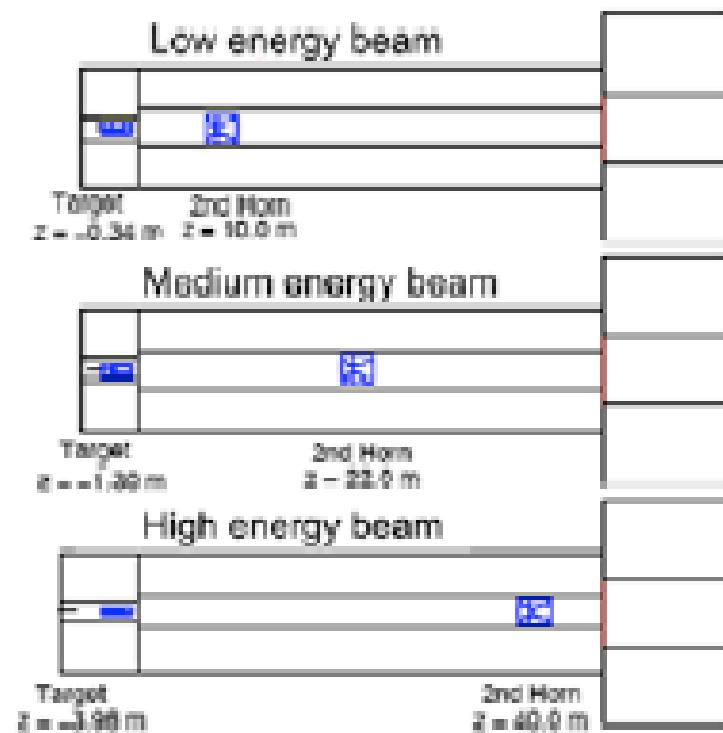
The Experimental Approach

- Depends on the “initial conditions”...
 - Existing accelerator ?
 - Proton/neutrino energies
 - Existing beamline ?
 - Beam direction
 - Desirable site ?
 - baseline
 - Underground?
 - On surface?
 - Existing Detector ?
 - Detection threshold/efficiency
 - Baseline
- 
- Fermilab :
50 - 120 GeV protons
NuMI beam :
neutrinos aimed down
At 57 mrad to Soudan,
Minnesota 700 km <L< 900km

The Experimental Approach

- Existing Conditions or Optimized Parameters (i.e.)
 - baseline (L)
 - initial beam composition ($\nu_\mu, \bar{\nu}_\mu, \nu_e$)
 - unoscillated energy spectrum (E)
 - detection efficiency
- ... will determine
 - # of events expected in the absence of oscillation
- Physics will determine
 - # of events (of each flavor)
 - *appearing and disappearing* due to **oscillation parameters, CP phase and matter effects**
 - ν and anti- ν rates differ due to **mass hierarchy and CP phase**
- Multiple measurements are needed to disentangle effects

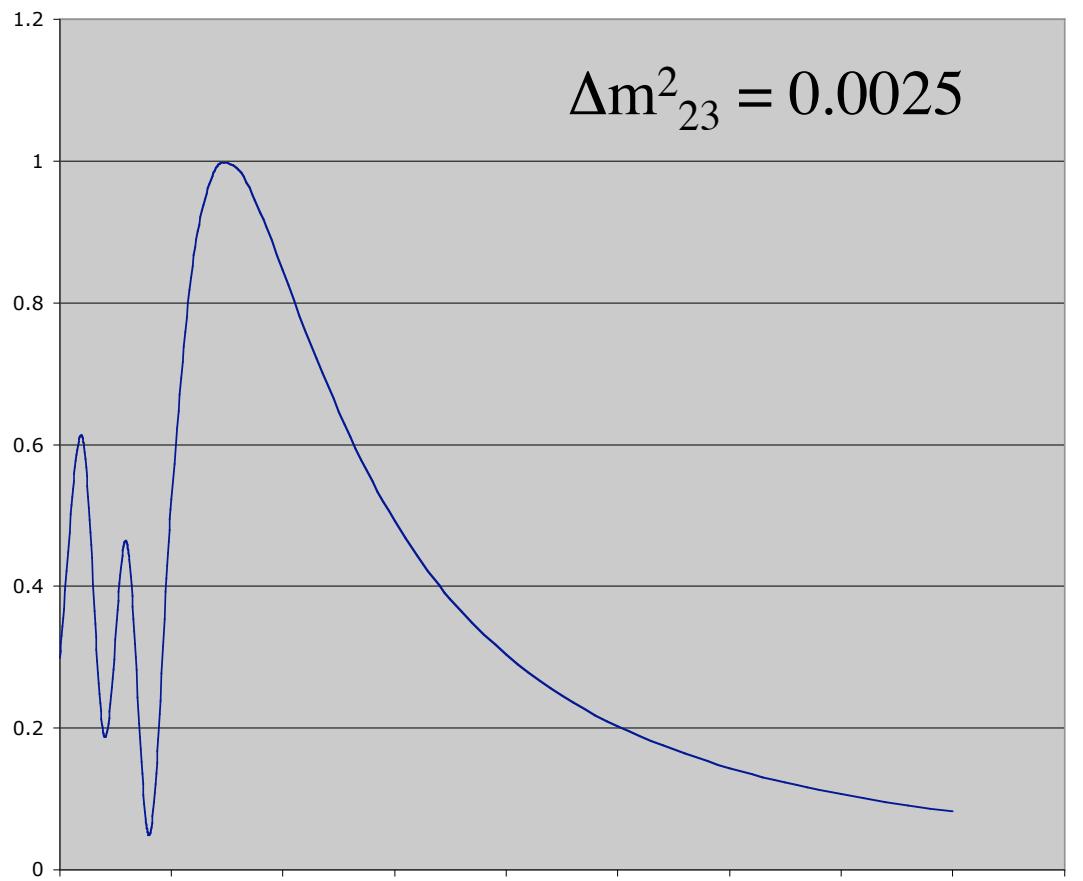
Variable Energy NuMI Beam



Oscillations at L = 735 km

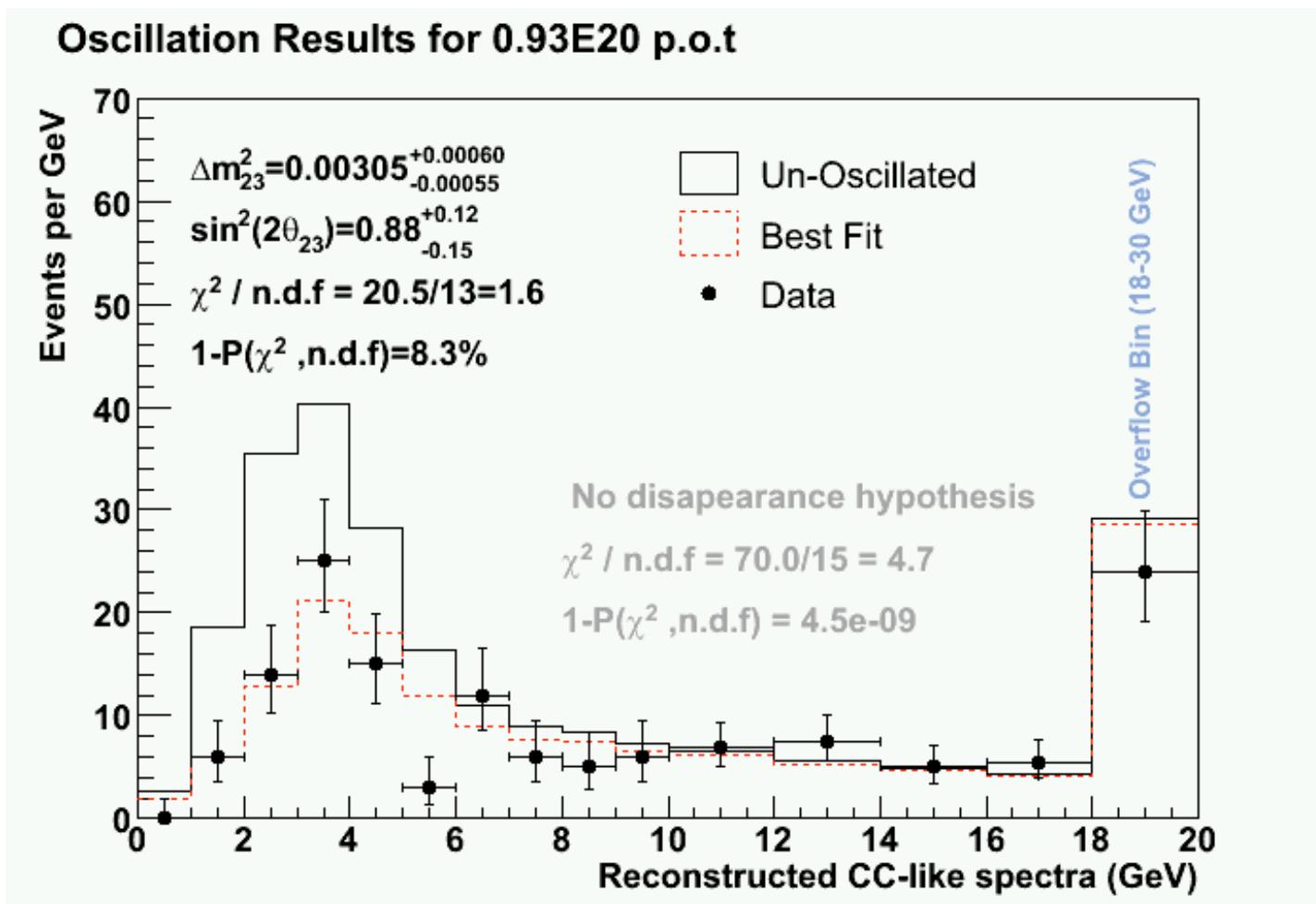
L determined by an existing Laboratory

- $\nu_\mu \rightarrow \nu_\tau$



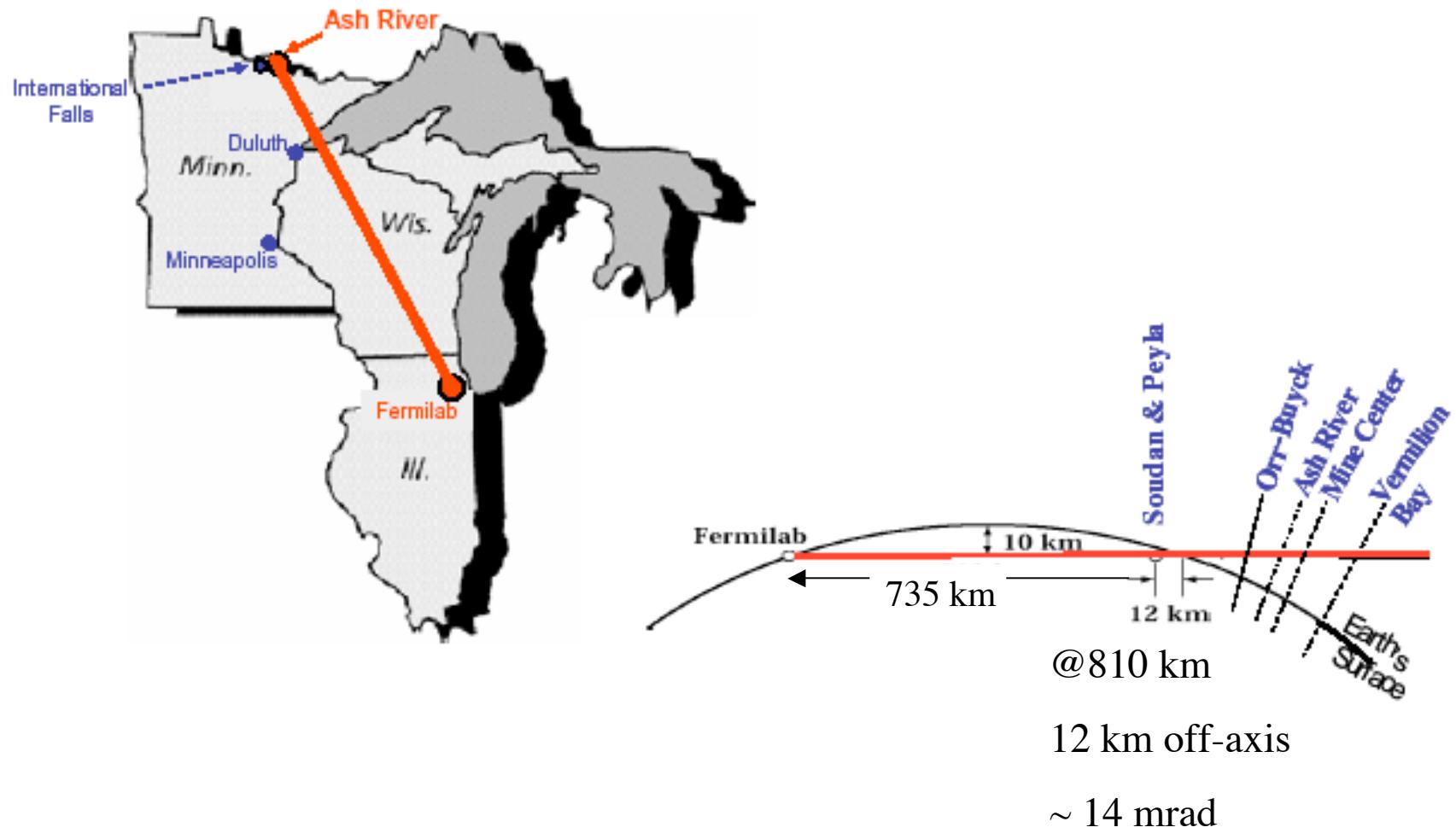
Presentation to NuSAG
May 20, 2006

Dominant $\nu_\mu \rightarrow \nu_\mu$ disappearance *a.k.a. MINOS*



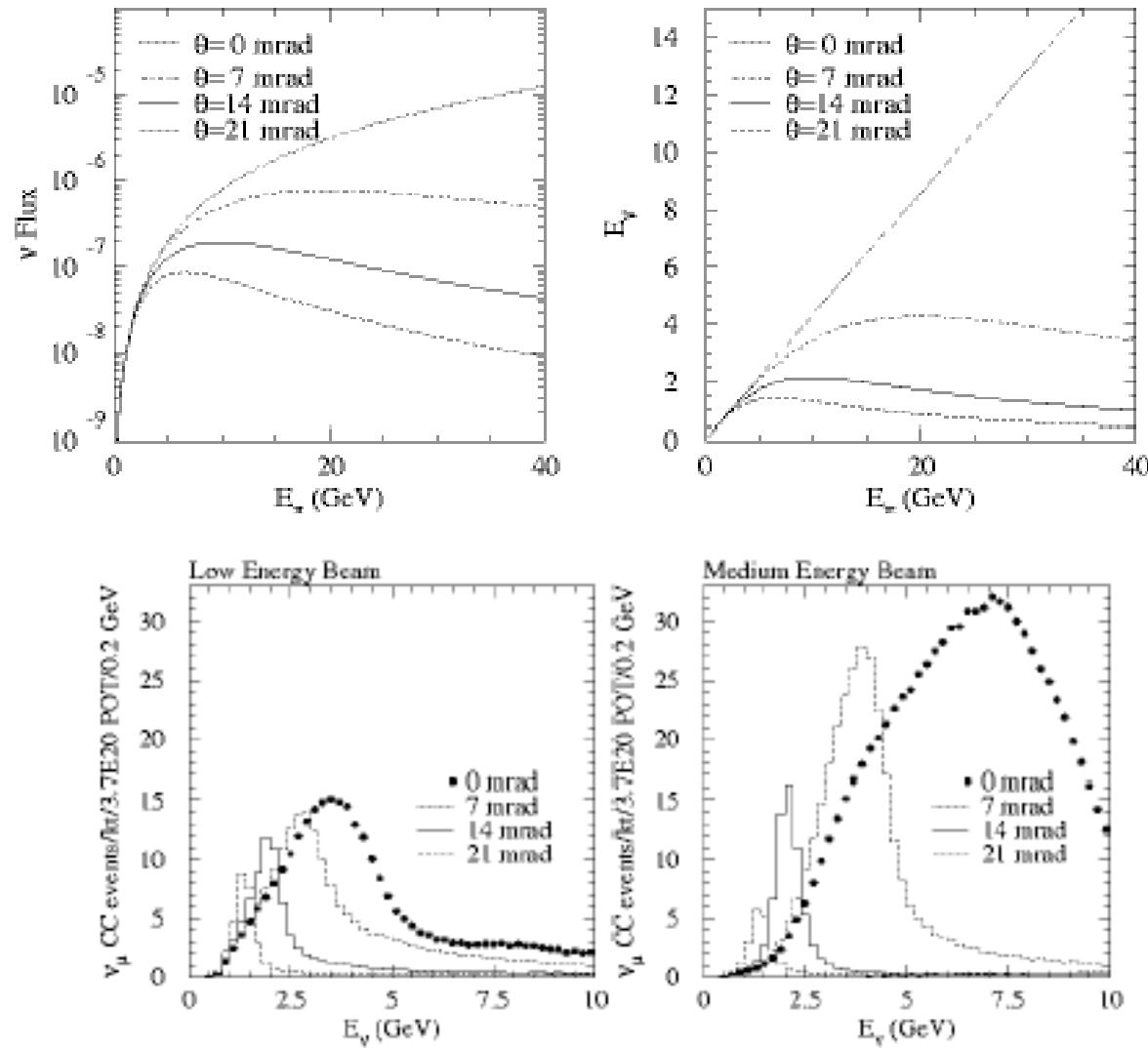
Presentation to NuSAG
May 20, 2006

The Experimental Approach : “NuMI North”



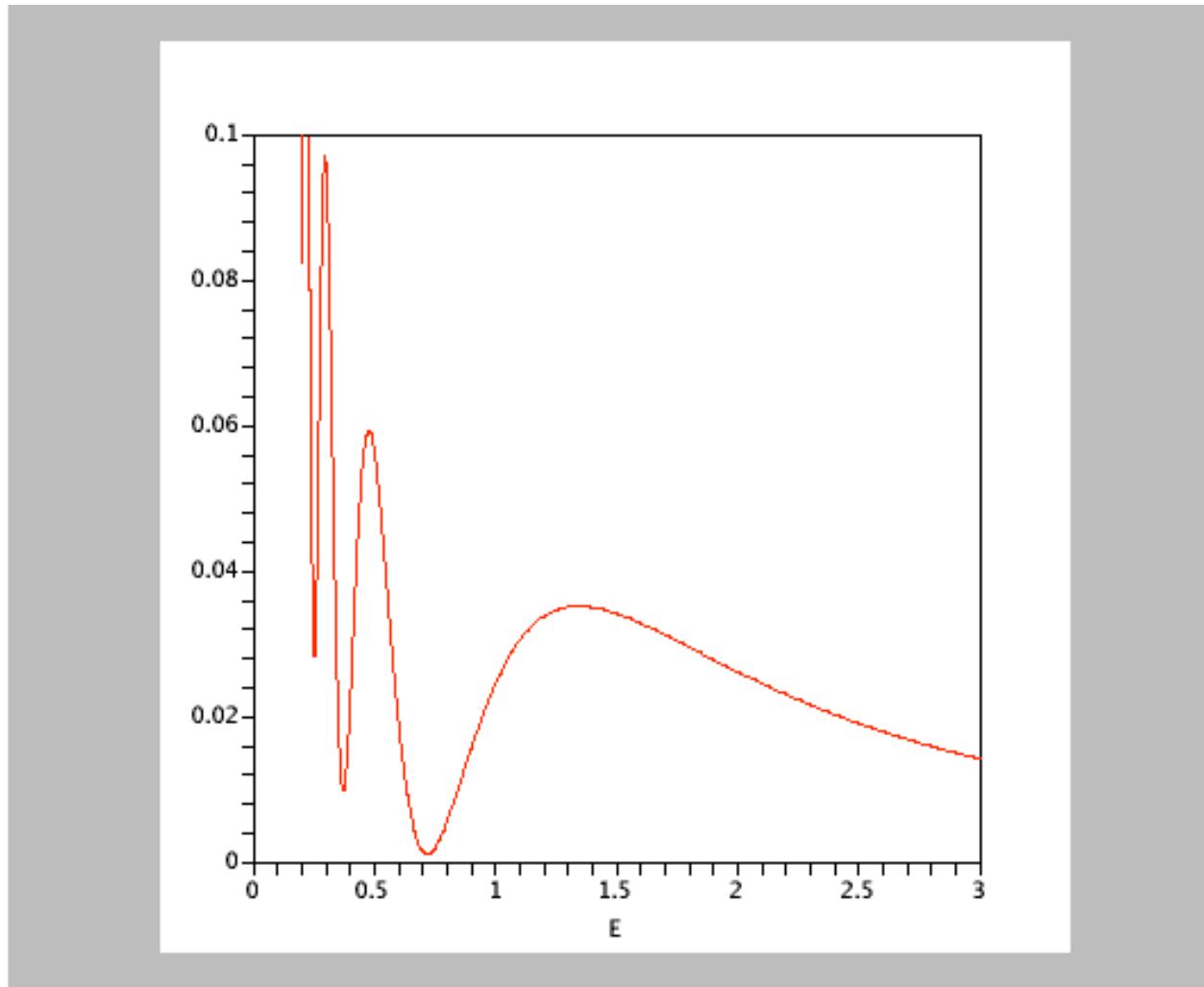
Presentation to NuSAG
May 20, 2006

The Off-axis concept



Presentation to NuSAG
May 20, 2006

Sub-dominant $\nu_\mu \rightarrow \nu_e$ appearance



Example :
 $L = 810 \text{ km}$
 $\Delta m^2_{23} = 0.0025$
 $\theta_{23} = 0.642$
 $\theta_{13} = 0.15$

Event Rates at the 1st Maximum

$$P_{vac} (\nu_\mu \rightarrow \nu_e) = \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \Delta_{atm},$$

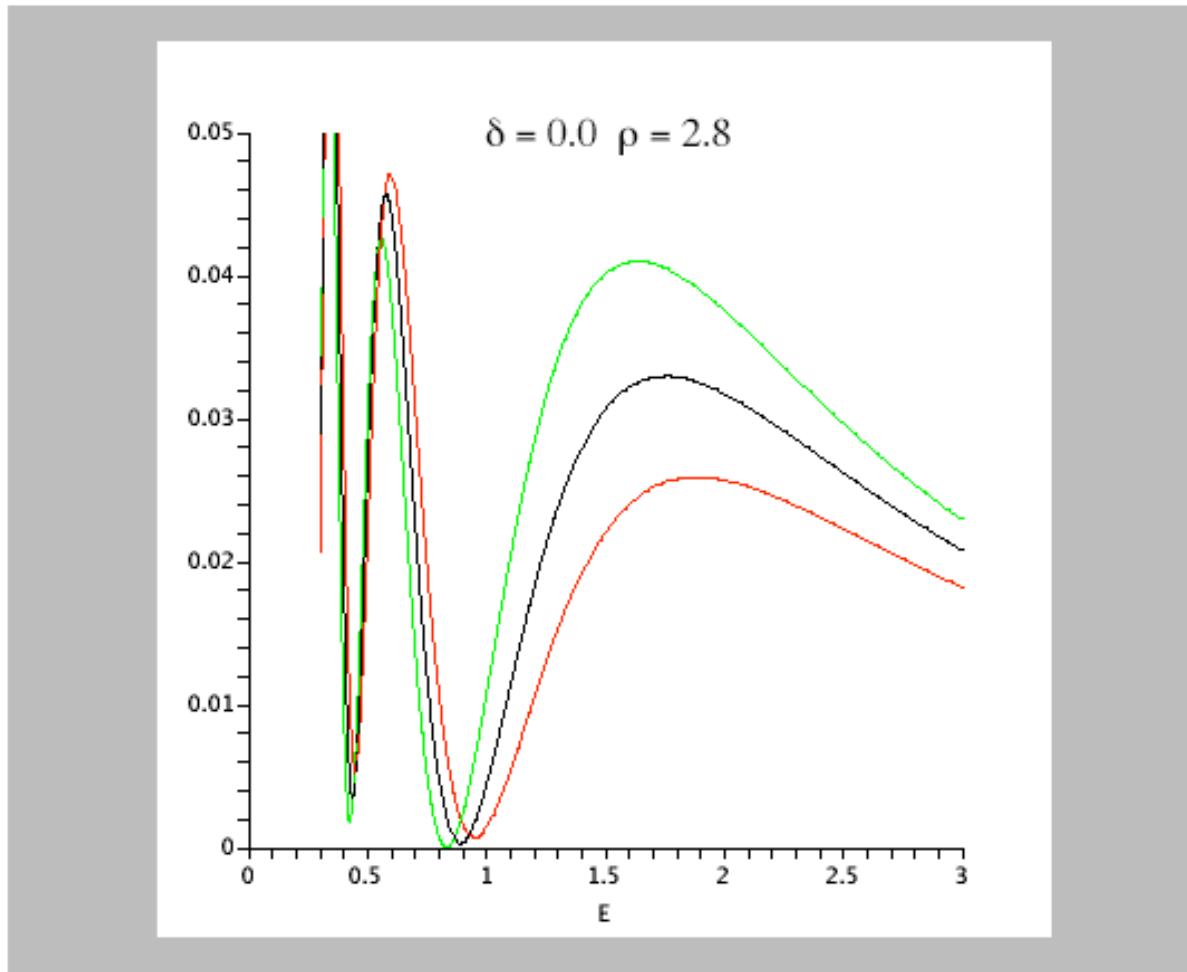
$$\Delta_{atm} \approx 1.27 \left(\frac{\Delta m_{32}^2 L}{E} \right),$$

- Determining θ_{13}
 - Input physics parameters :
 - $\sin^2 \theta_{23}$
 - $\Delta m_{atm}^2 \sim \Delta m_{32}^2$
 - Neutrino cross sections
 - Input experimental parameters
 - Protons per year
 - Neutrino spectrum
 - Detector Location ($L, \Delta x$)
 - Detector fiducial mass, efficiency

Presentation to NuSAG

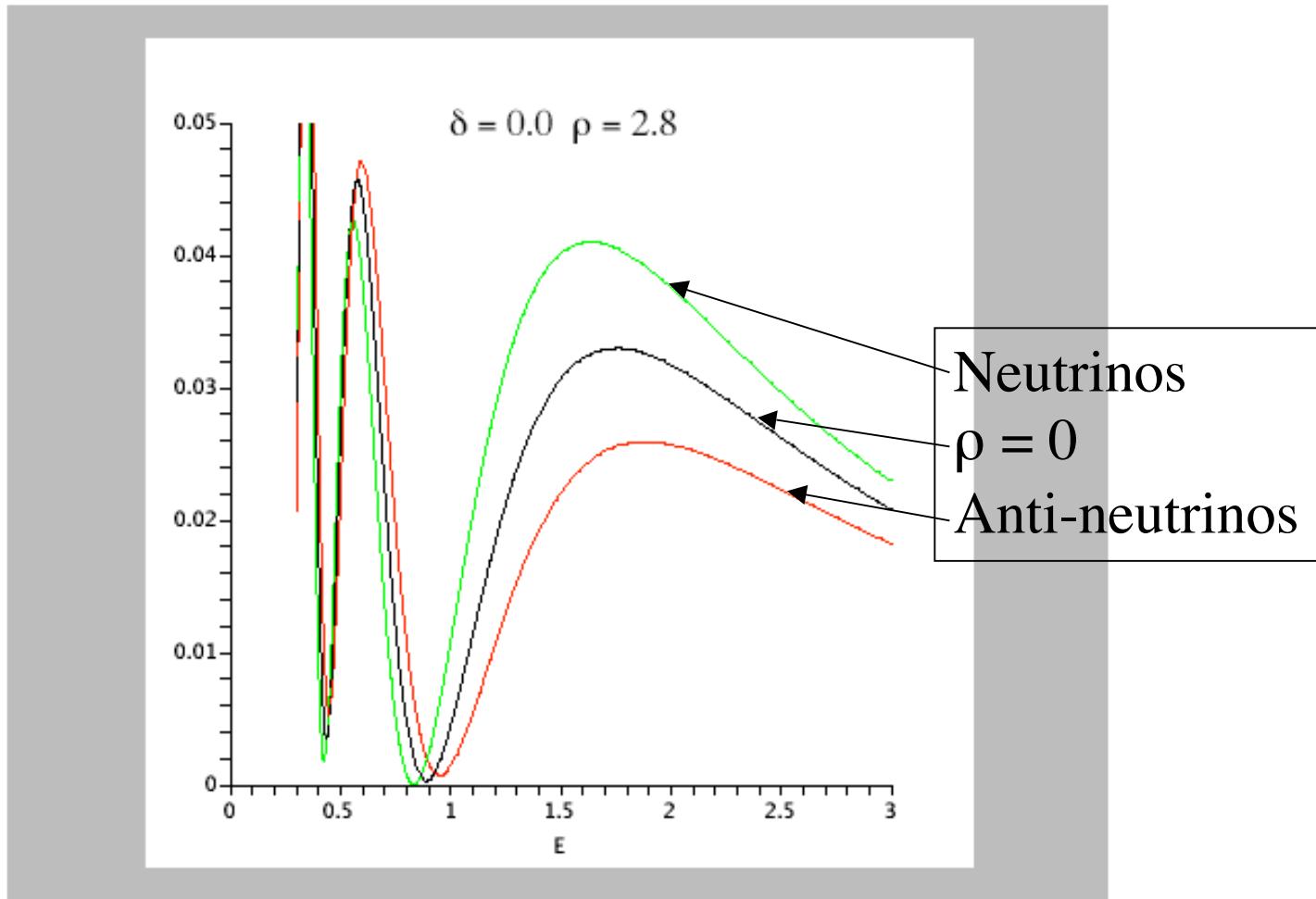
May 20, 2006

Oscillations affected by matter

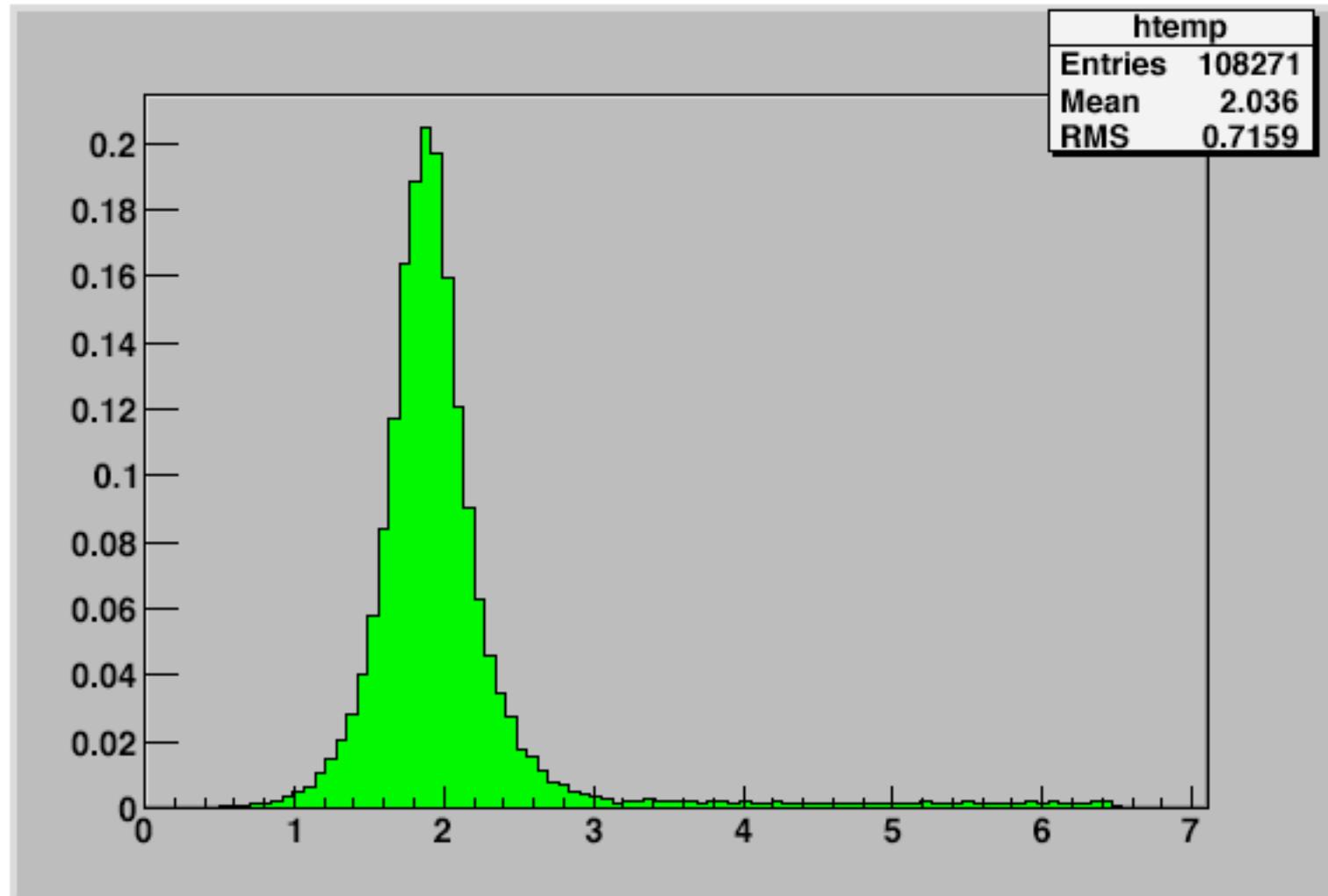


Presentation to NuSAG
May 20, 2006

Oscillations affected by matter

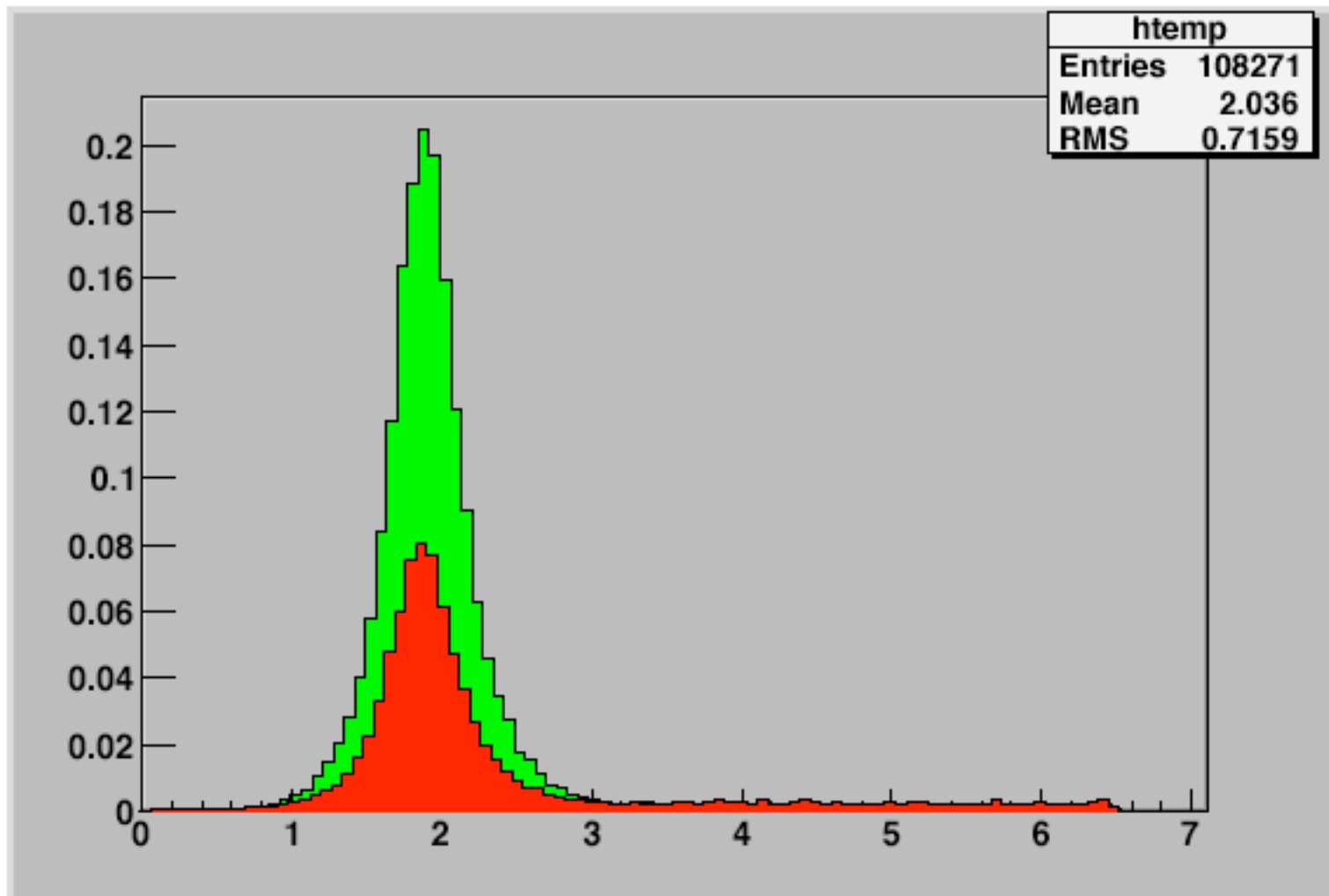


Create a beam : 810 km, 14 mr off xis



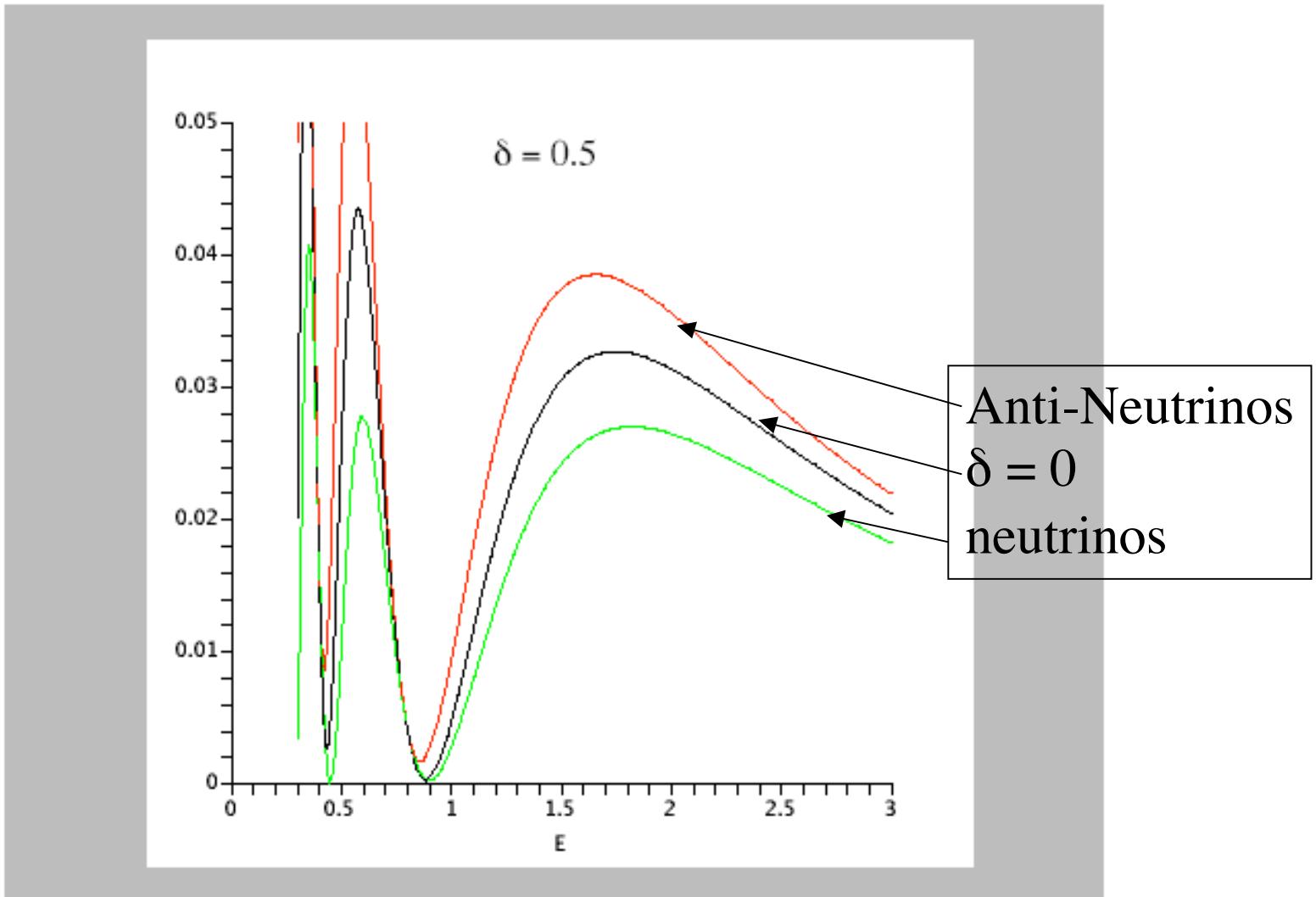
Presentation to NuSAG
May 20, 2006

Change horn currents (anti-neutrinos)



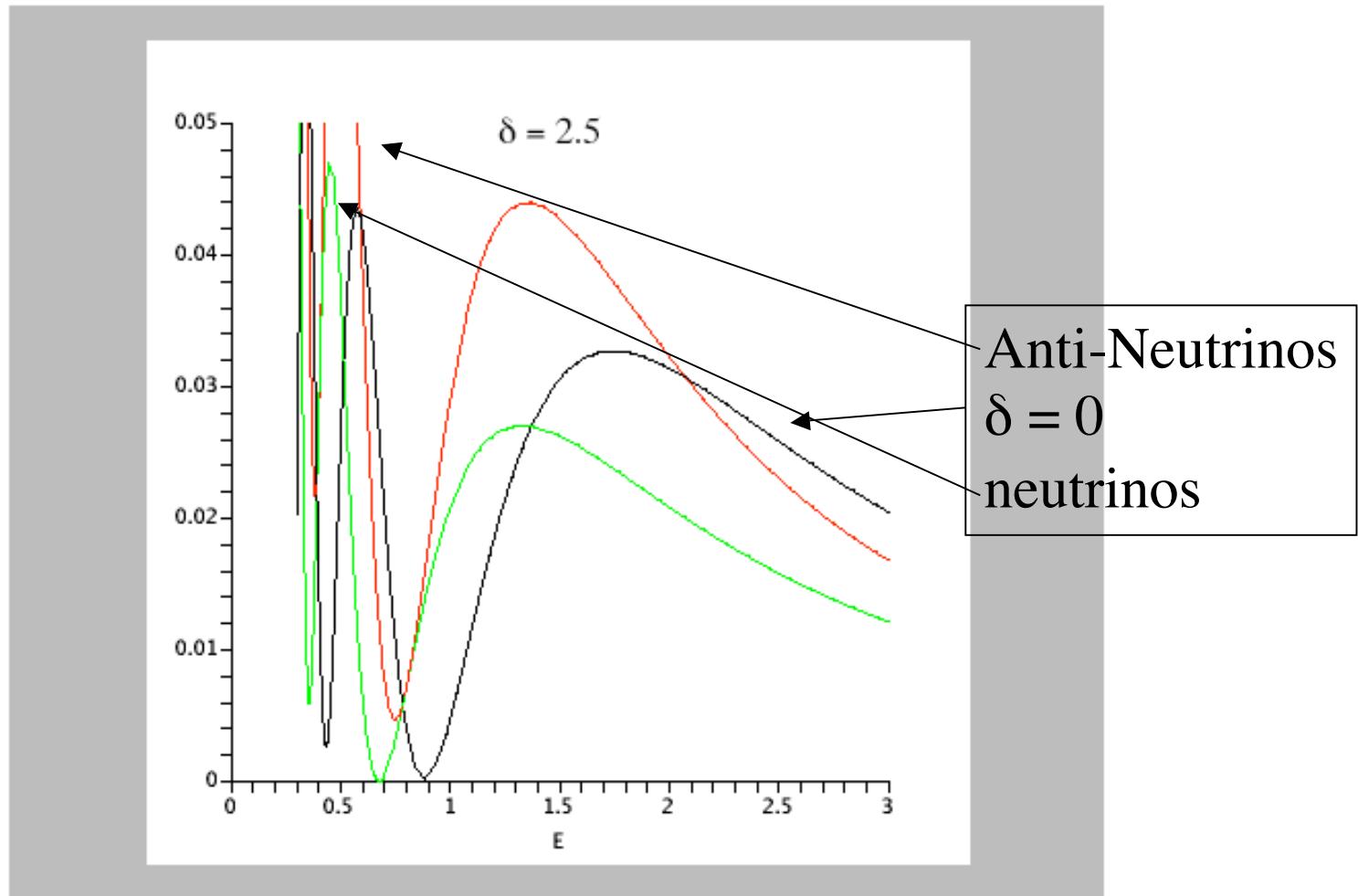
Presentation to NuSAG
May 20, 2006

But life isn't simple...

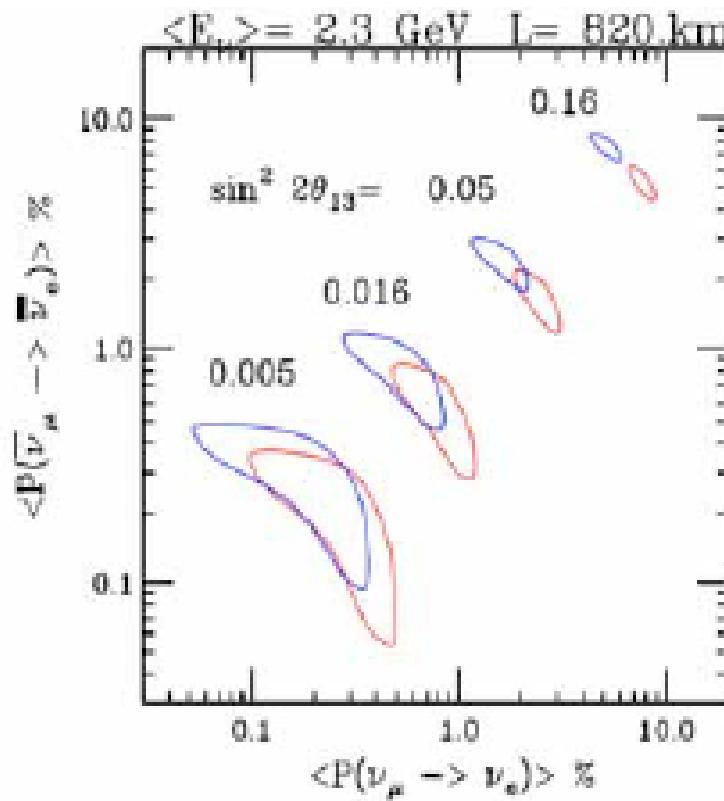
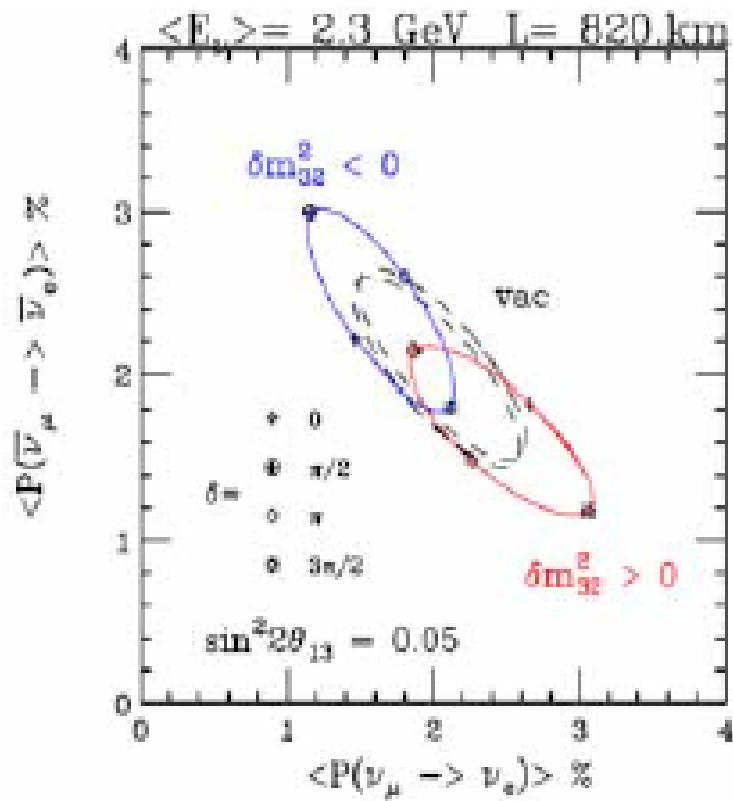


Presentation to NuSAG
May 20, 2006

And we have no idea what δ is...

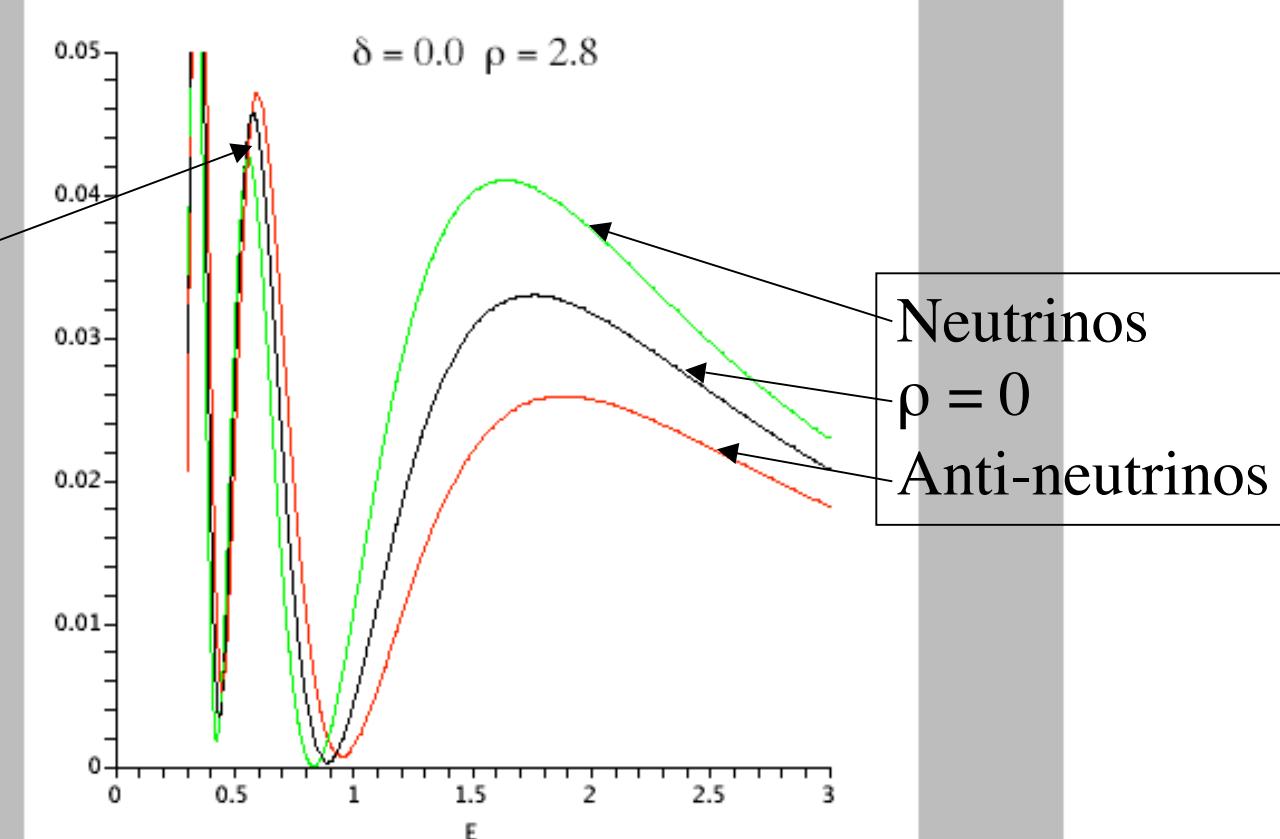


Measurements ?

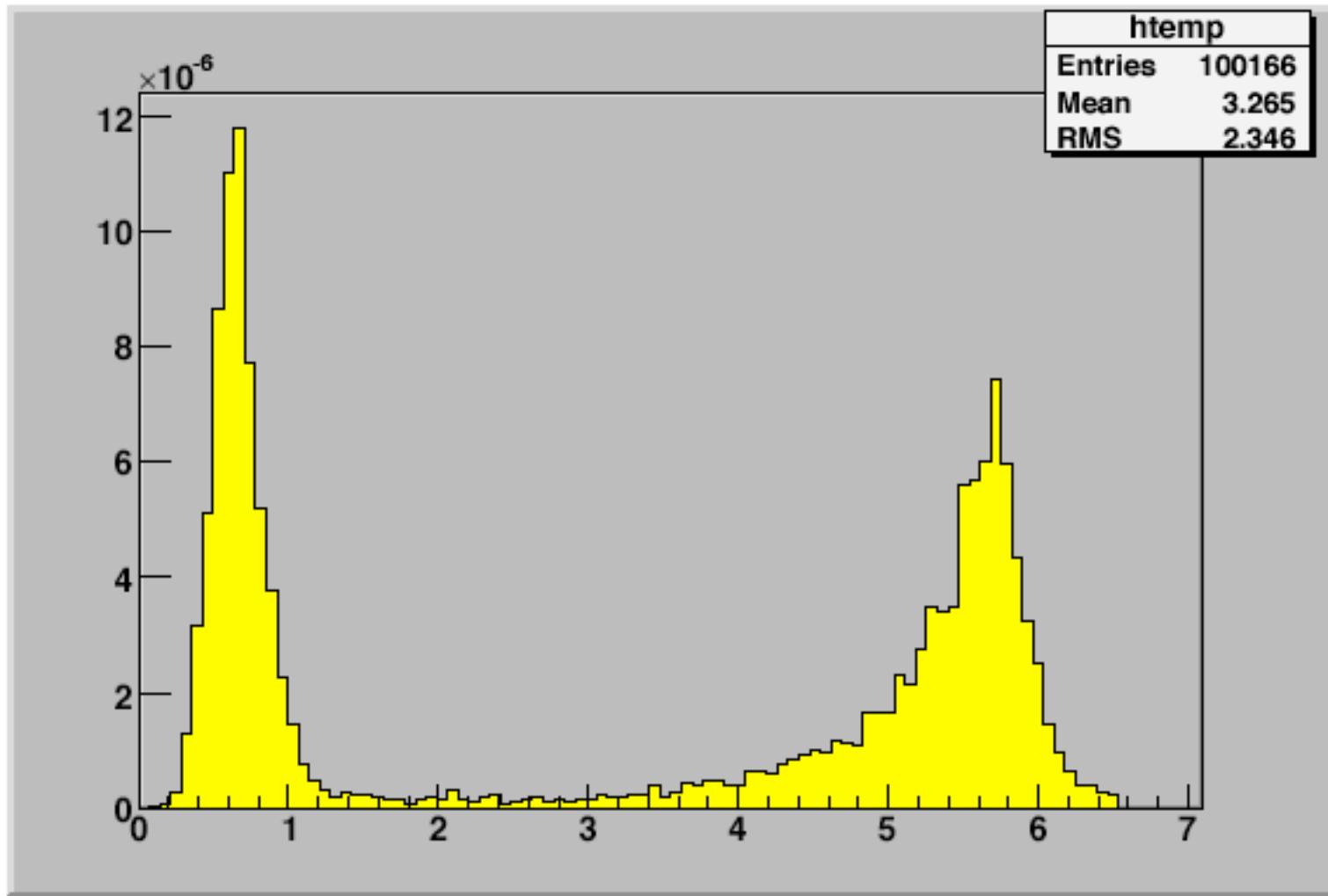


Oscillations affected by matter

But not so
much at the
2nd maximum

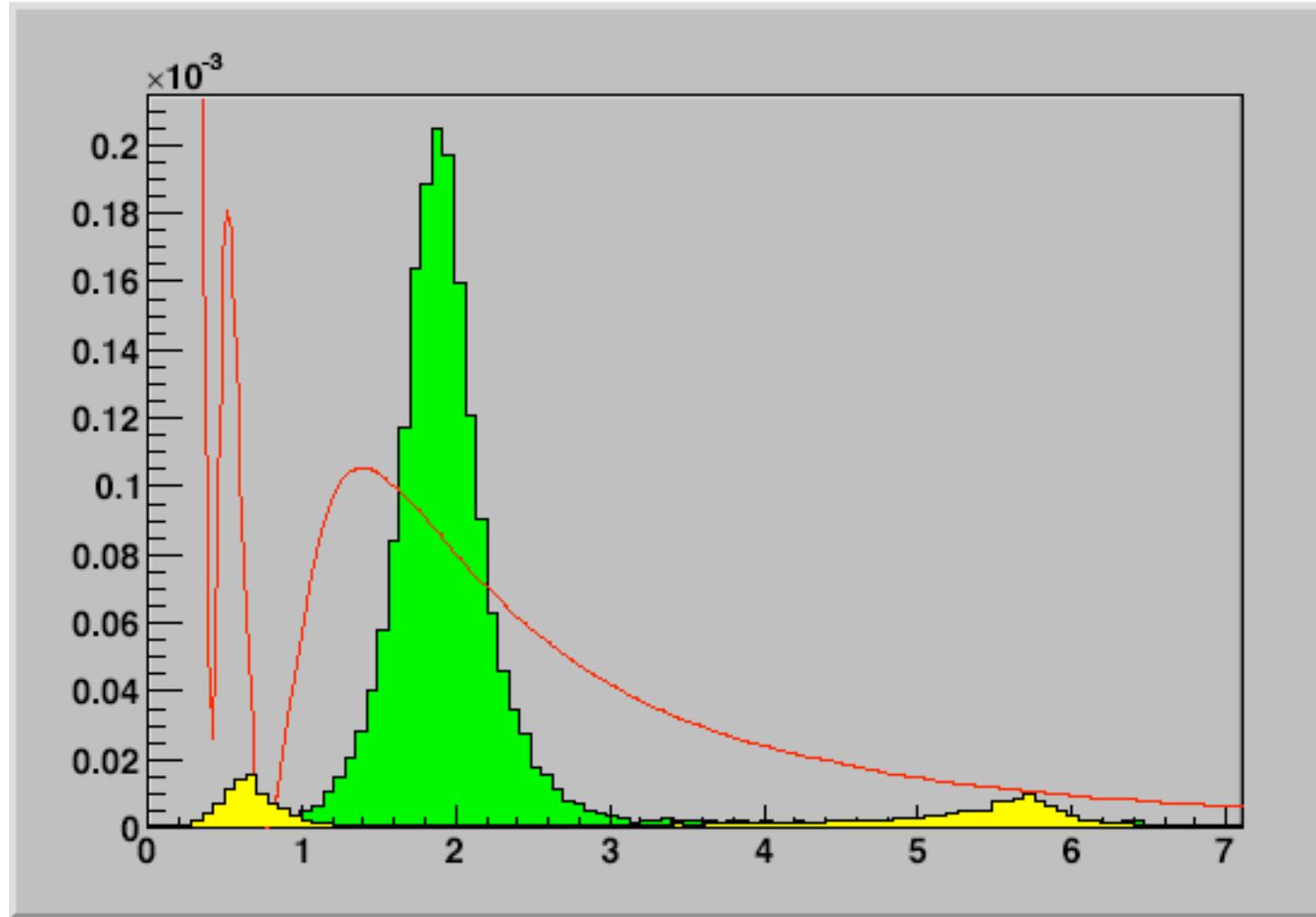


NuMI Off-axis - 40 mrad (32 km)



Presentation to NuSAG
May 20, 2006

$\nu_\mu \rightarrow \nu_e$ appearance



Presentation to NuSAG
May 20, 2006

Example : $L = 810$ km
 $\Delta m^2_{23} = 0.0025$

Event Rates

$L = 810 \text{ km}$

12 km OA

20 kt fiducial

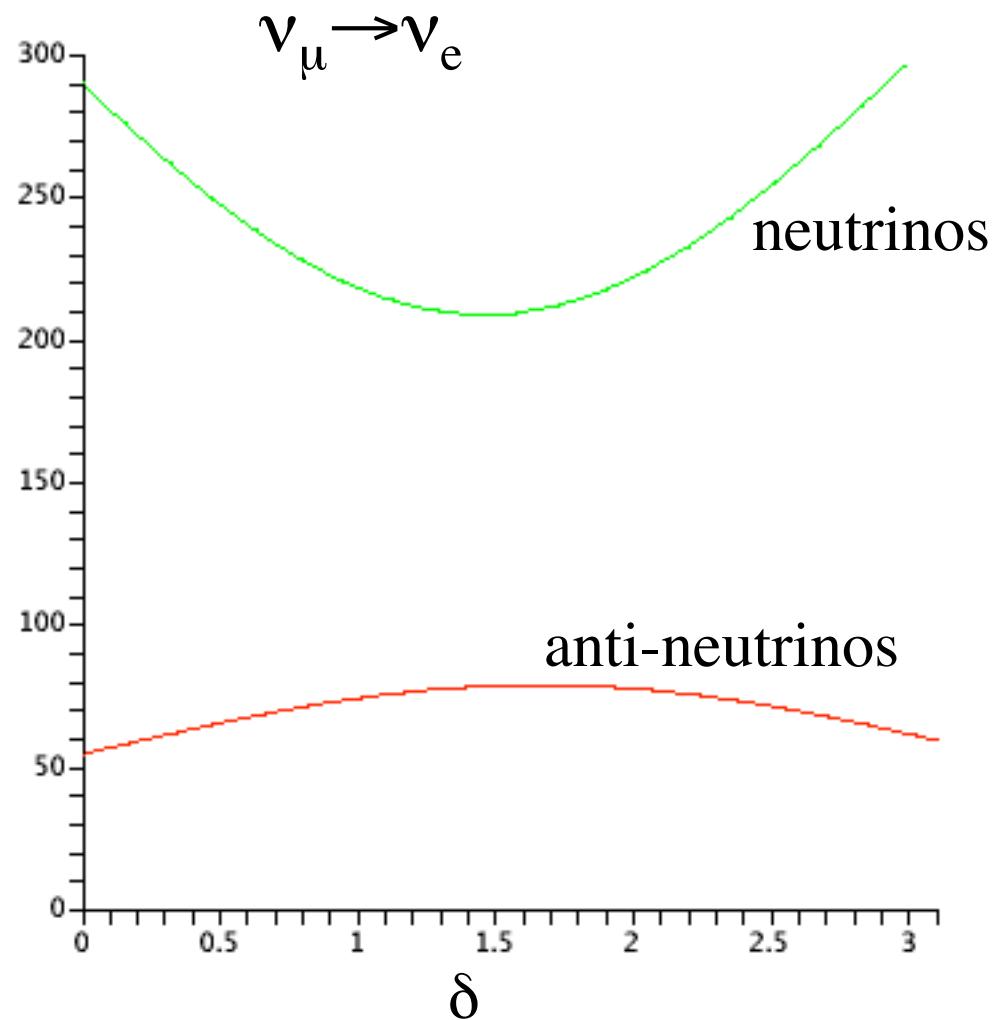
20e20 pot ($\sim 3 \text{ yrs}$)

(for each mode)

$\Delta m^2_{23} = 0.003$

$\theta_{13} = 0.15$

Normal hierarchy



Event Rates

$L = 810 \text{ km}$

40 km OA

20 kt fiducial

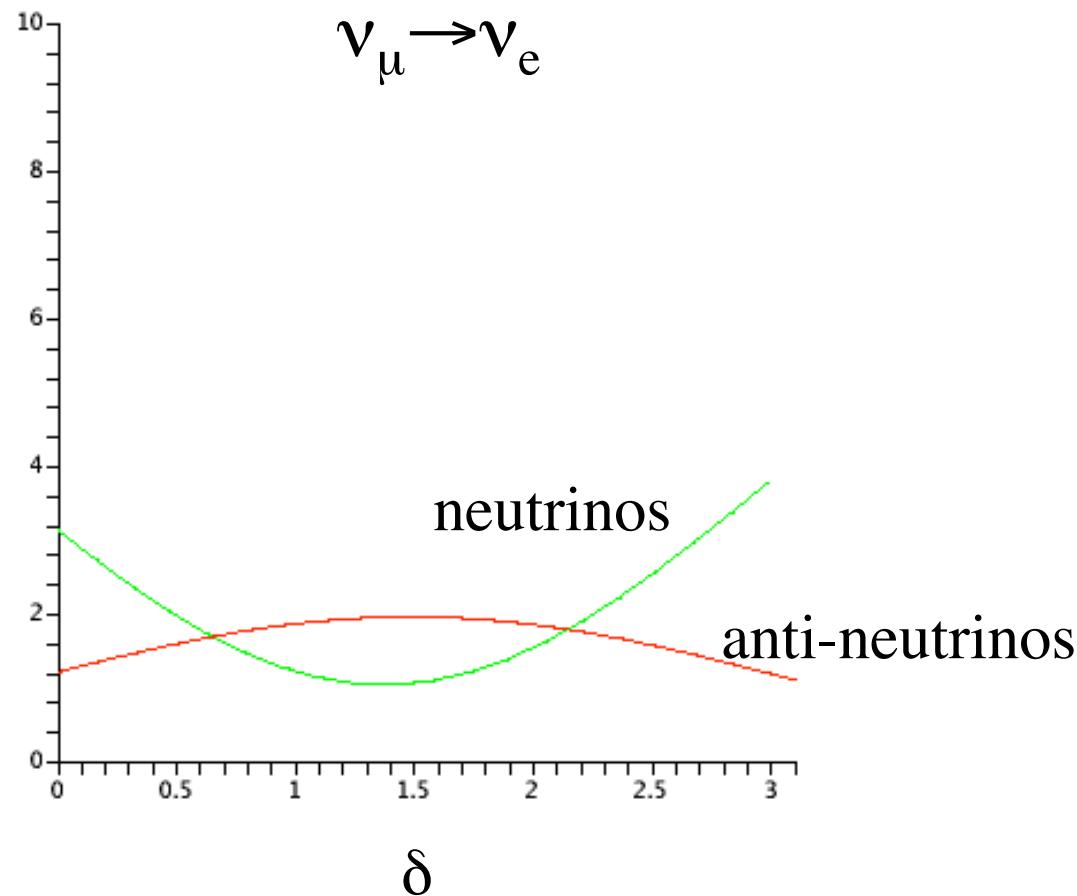
20e20 pot ($\sim 3 \text{ yrs}$)

(for each mode)

$\Delta m^2_{23} = 0.003$

$\theta_{13} = 0.15$

Normal hierarchy



Backgrounds

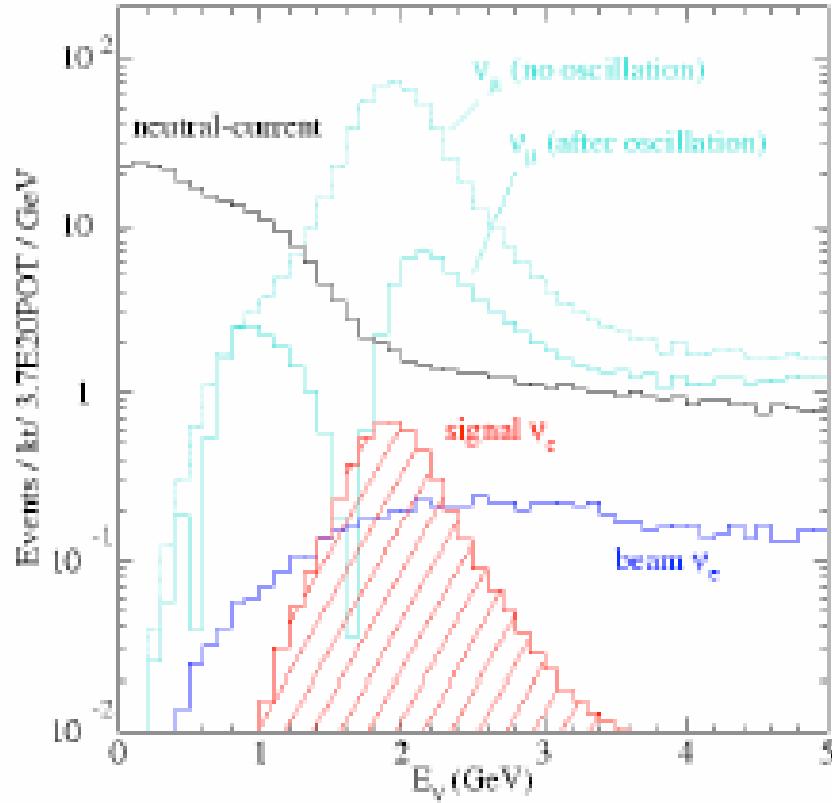
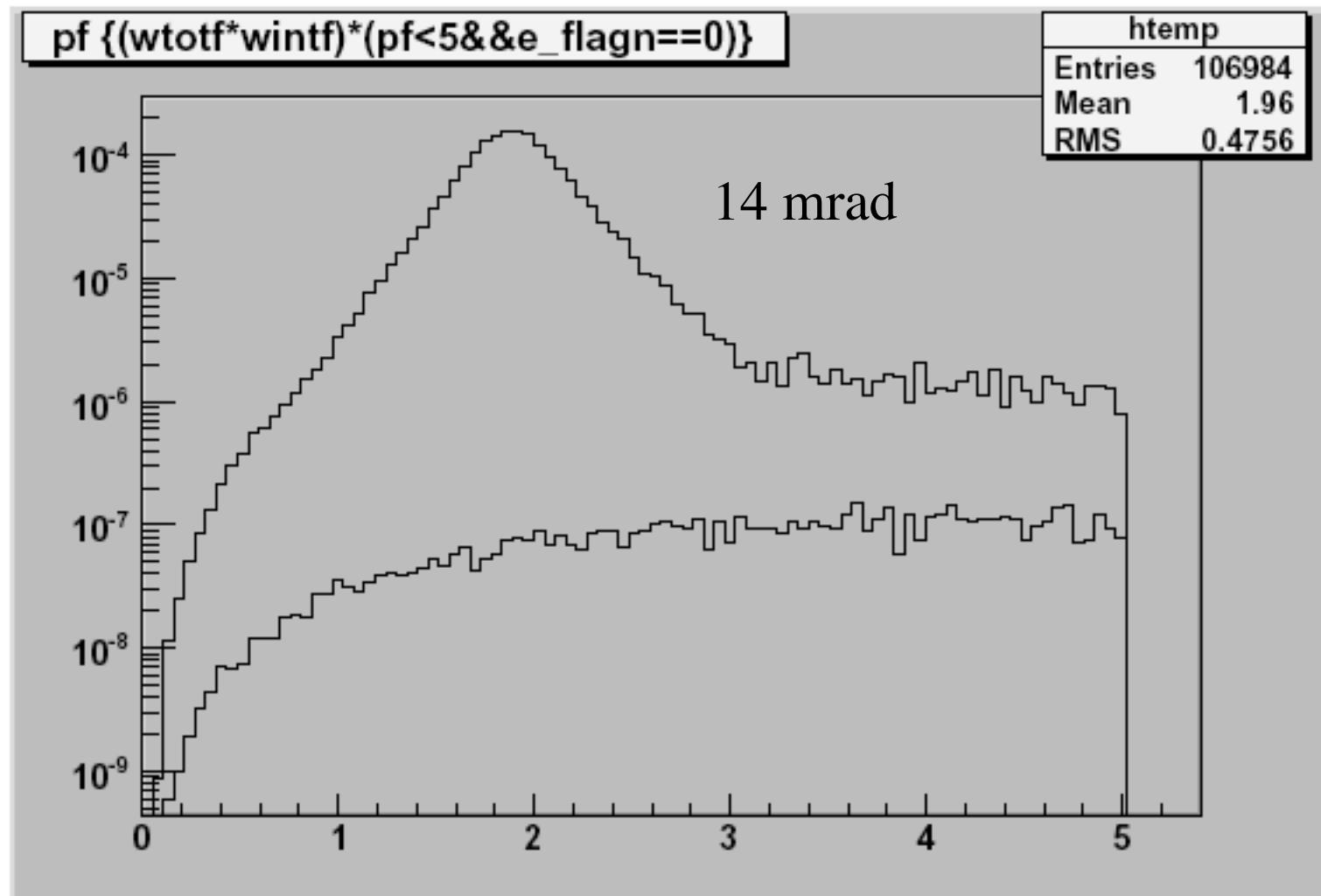


Fig. 2.8: Simulated energy distributions for the v_e oscillation signal, intrinsic beam v_e events, neutral-current events and v_μ charged-current events with and without oscillations. The simulation used $\Delta m^2_{32} = 2.5 \times 10^{-3} \text{ eV}^2$, $\sin^2(2\theta_{23}) = 1.0$, and $\sin^2(2\theta_{13}) = 0.04$. An off-axis distance of 12 km at 810 km was assumed.

Backgrounds (cont)



Presentation to NuSAG
May 20, 2006

Conclusions (1)

- An Off-axis neutrino beam is a powerful tool for measuring oscillation parameters and the neutrino mass hierarchy
- The success of experiments using this technique depend heavily on the ability of the proton source to deliver lots of protons
- In the best of worlds $\sin^2 2\theta_{13} \sim 0.1$ the experiments are difficult and take a long time to proceed through the list of questions
- If nature is cruel, $\sin^2 2\theta_{13} \sim 0.01$, we will most likely need multi-megawatt proton beams aimed at megaton detectors.

Conclusions (2)

- FMI Operations (January 2005 -February 2006) have delivered $1.4\text{e}20$ protons to the NuMI target ($1.3\text{e}20$ in the Low energy configuration)
- MINOS experiment is counting on getting $\sim 3\text{e}20$ protons/year for 2007 - 2009
- The NOvA experiment is planning on $\sim > 6.5\text{e}20/\text{year}$ (700 kw) starting in 2011 (see Alberto's talk)

Conclusions (3)

- An off-axis experiment at the 2nd oscillation maximum will require :
 - increased protons (> 1 MW) ?
 - more mass (50 - 100 kton) ?
 - higher efficiency ($\sim 80\%$) ?
 - longer running time (10 yrs) ?
- Measuring the neutrino sector CP phase
 - priceless

On-going/Upcoming Work

- Complete event rate analysis for 1st and 2nd maximum at $L \sim 810$ (12 and 32 km off-axis)
- Study $L \sim 250$ km sensitivity
- Study sensitivity at longer baselines (1500, 2500 km)
- Validate background estimates for beam
- Compare to wide band beam scenarios

A final comment

- In addition to the experiments being difficult from the signal to background point of view, collecting a handful events/year for many years will require some new thinking about how to manage and retain experimental collaborations
- Expanding the capability of the detectors to have more physics reach may be a necessary investment